

ADVANCED CYSYLLTIADAU CONNECTIONS PELLACH



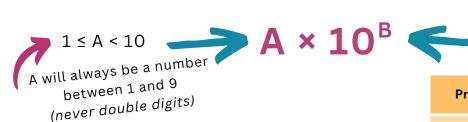
SESSION 2: PERFORMING LIKE A PHYSICIST

The aim of this tutorial is to consider how we think about expressing and analysing values and measurements. More specifically, using standard form to express values, and using accuracy and precision to scrutinise them.

WHAT IS STANDARD FORM?

AND WHY DO WE USE IT?

In physics, we deal with a lot of very large numbers and a lot of very small numbers. Writing these numbers out in full is time-consuming, makes calculations unnecessarily difficult, and it's not immediately obvious how big or small these values are. Instead, we use standard form:



Using this form to express orders of magnitude is so useful that physicists do it all the time! If you've ever used millimetres (mm), kilograms (kg), or megabytes (MB) then you might have done this without even realising. The prefixes m, k, and M simply correspond to a standard form (also known as a "power of 10").

	B is called the
	order of magnitude

Prefix	Symbol	Standard Form
Tera	Т	10 ¹²
Giga	G	10 ⁹
Mega	М	10 ⁶
Kilo	k	10 ³
Milli	m	10 ⁻³
Micro	μ	10 ⁻⁶
Nano	n	10 ⁻⁹
Pico	Р	10 ⁻¹²

EXAMPLE: GRAVITY

Let's say we wanted to find the gravitational force between the Earth and the Moon according to this equation:

$$F=rac{Gm_1m_2}{r^2}$$
 where

 $G=0.0000000000667~{
m Nm^2/kg^2}$

 $r = 385000000 \,\mathrm{m}$

In standard form, we would write these as:

 $G = 6.67 \times 10^{-11} \, \mathrm{Nm^2/kg^2}$ $m_1 = 5.97 \times 10^{24} \, \mathrm{kg}$ $m_2 = 7.34 \times 10^{22} \, \mathrm{kg}$ $r = 3.85 \times 10^8 \text{m}$

MUCH EASIER TO READ!



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CALCULATIONS USING STANDARD FORM

We use these rules when multiplying or dividing values in standard form:

$$A^x imes B^y = (A imes B)^{x+y}$$
 and $rac{A^x}{B^y} = (A imes B)^{x-y}$

$$\frac{A^x}{B^y} = (A \times B)^{x-y}$$

EXAMPLE 1

$$\left(8\times10^7\right) imes\left(4\times10^3\right)$$

$$=(8\times4) imes(10^7 imes10^3)$$

$$=32 imes10^{7+3}$$
 Apply the Rule

$$=32 imes10^{10} \ =3.2 imes10^{11}$$

EXAMPLE 2

$$egin{aligned} \left(8 imes 10^7
ight) \div \left(4 imes 10^3
ight) \ &= \left(8 \div 4
ight) imes \left(10^7 \div 10^3
ight) \end{aligned}$$

$$=2\times10^{7-3}$$

 $=2 imes10^{7-3}$ Apply the Rule

$$=2 imes10^4$$

WE DON'T NEED TO CONVERT, BECAUSE 2 IS ALREADY BETWEEN 1 AND 10

TASK 1: CALCULATION PRACTICE

15-20 MINS

Now try some calculations yourself:

1.
$$(3 \times 10^4) \times (8 \times 10^2) = ?$$

$$\frac{2. \ \ \frac{\left(8.4 \times 10^6\right)}{\left(2.1 \times 10^2\right)} = ?$$

3.
$$(9 \times 10^2) \times \frac{(5 \times 10^9)}{(2.5 \times 10^3)} = ?$$

Can you prove these harder calculations?

4.
$$(5 \times 10^6) \times \frac{2.4 \times 10^2}{1.2 \times 10^8} = 10$$

$$5. \ \, \frac{10^3 \times 10^{-2} \times 10^7}{10^{-5} \times 10^4} = 10^9$$

TASK 2: GRAVITY

10 MINS Can you use the rules above and the values on the previous page calculate the standard form of F?

$$F=rac{Gm_1m_2}{r^2}$$

Challenge: try working out the order of magnitude first without using a calculator!





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ACCURACY & PRECISION

The precision of a value is a lot clearer when we put really big or really small numbers in standard form.

But what exactly do we mean by **precision?**

And how is it different to accuracy in the context of physics?

THIS ISN'T ALWAYS SOMETHING THAT WE KNOW!

ACCURACY How close a measurement is to the true value



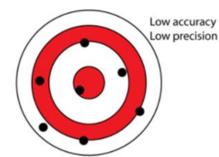
WHAT WE WANT AS PHYSICISTS:

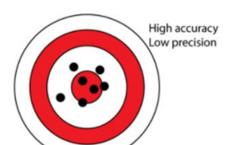


HIGH ACCURACY

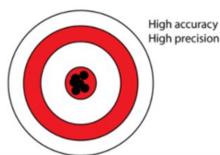


HIGH **PRECISION**









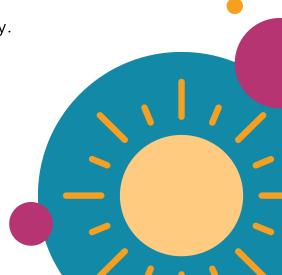
TASK 3: ACCURACY VS PRECISION

5 MINS

A person decides to measure what time sunset is each day. What they do:

- 1. They use a radio-controlled clock
- 2. They take a reading to the nearest 100 ms as soon as sunlight enters their room in the morning
- 3. The horizon to the east (direction of sunrise) is obscured by a few trees and buildings

How would you describe these readings in terms of accuracy and precision?





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TASK 4: RECORDING DATA

5 MINS

What you have learned so far is very important when it comes to how we record data. The following table contains many errors.

Can you think of at least 3 ways this table can be improved?

HINT: some errors relate to what you've learnt this session, and some don't!

(V)	Т	RC (Ω)	KIN (Ω)
6.28	293.2	227.4	13100
6.088	286.4	223	16440
5.8	276.7	217.2	22600
5.1	252.7	201	56270
5	249.2	198.9	64660
4.35	226.9	184.4	175900
4.0	216.9	177.7	284100
3.9119	211.7	174	372000
3.5	200.4	165.8	704000
3.0809	183.2	153.4	2055000
2.245	161.2	135.6	10600000
1.952	144.4	121.2	49400000
1.6	134.4	110.1	OL

OPTIONAL EXTRAS

Can you find the value of E in the following equations? Try **not** to use a calculator!

If you need to use a calculator, try to find the order of magnitude of E without one to check that your calculations were correct.

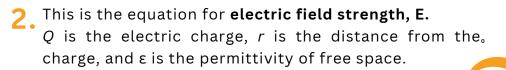
$$E=mc^2$$

1. Conservation of energy tells us that energy can't be created or destroyed. But it can be converted into other forms. Einstein discovered that it is even equivalent to matter!

$$m_e=$$
 9.11 × 10 $^{ ext{-}31}$ kg $c=$ 2.99 × 10 $^{ ext{8}}$ m/s

What is the **energy of** electron using these values?

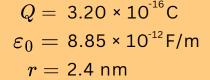
we are ignoring kinetic and potential energy



What is the **electric field strength** 2.4 nm away from an alpha particle using these values?









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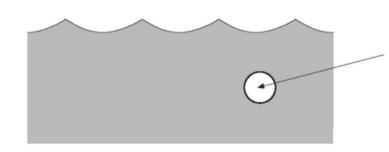
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EXAM PRACTICE

All exam questions in Advanced Connections are taken from WJEC A-level Physics papers!

Have a go at some of these exam questions:

A bubble of krypton (Kr) gas is formed underwater. The bubble has a volume of $2.14 \times 10^{-6} \,\mathrm{m}^3$ and contains 1.70×10^{20} krypton molecules each of mass $1.39 \times 10^{-25} \,\mathrm{kg}$. The pressure of the gas inside the bubble is $300 \,\mathrm{kPa}$.



Kr bubble volume = $2.14 \times 10^{-6} \,\mathrm{m}^3$ pressure = $300 \,\mathrm{kPa}$ No. of molecules = 1.70×10^{20} mass of a Kr molecule = $1.39 \times 10^{-25} \,\mathrm{kg}$

- (a) Calculate:
 - (i) the rms speed of the Kr molecules in the bubble;

[3]

[2]

[3]

(ii) the relative molecular mass of Kr;

- (iii) the temperature of the gas inside the bubble.
- Information about the Earth, Jupiter and the Sun is given in the table:

	Radius / m	Mass / kg	Distance from the Sun / m
Earth	6.37×10^{6}	5.97×10^{24}	1.50×10^{11}
Jupiter	6.99 × 10 ⁷	1.90×10^{27}	7.79×10^{11}
Sun	6.96 × 10 ⁸	1.99×10^{30}	_

- (a) Determine the following at the position of the Earth.
 - (i) The gravitational field strength of the Sun.

[2]

(ii) The gravitational potential due to the Sun.

[2]

(b) Show clearly that the centre of mass of the Sun-Jupiter system lies outside the Sun. [3]